The Chandra X-ray Observatory

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The Chandra X-ray Observatory

Launched 5 years ago 23 July 1999

A revolution in X-ray astronomy and astronomy in general

What can Chandra do?

Some science examples

Some comments on X-ray data
When we look up at the night sky
we see it filled with stars

But,
Outside the narrow range of colors
our eyes are sensitive to,
something quite different dominates
the night sky…

X-rays: A fundamental difference
Powerful sources of X-rays

A power source entirely different from the nuclear fusion that drives the Sun and stars

...and much more efficient

X-ray map of the whole sky:

100,000 `sources`

Rosat All Sky Survey (MPE)
We are now in the era of multiwaveband astronomy.
Sources of X-rays

- Shocked plasma
- Relativistic synchrotron plasma (ang. mom. + B field)
- Energy release from gravity (accretion power)

Explosions: Supernovae and their remnants

Particles moving near the speed of light in magnetic fields

Matter falling into deep gravitational wells

In the optical, we see mostly reprocessed fusion energy
In X-rays, we see mostly accreting sources: energy from gravity!
X-rays in a galaxy

- X-ray binary systems (accretion onto compact objects)
- Nuclear X-ray source (accretion onto massive BH)
- Supernova remnants
- Gas heated by stellar processes and galaxy interactions
40 Years of X-ray Astronomy: 1 billion times more sensitive

- 1962: Good for 1 (one) Nobel Prize
- 1962: Sco X-1: the brightest source of X-rays in the sky
- 1978: Good enough to ID bright galaxies
- 1999: Sco X-1: the brightest source of X-rays in the sky
- 2001: Chandra
- 2001: Distant galaxy 100,000 times fainter than NGC3783

Resolution is the key
Chandra takes X-ray Astronomy from its ‘Galileo’ era to its ‘Hubble’ era in a single leap

X-ray astronomy took just 40 years to match 400 years of optical astronomy
What is Chandra?

Chandra is the greatest X-ray Observatory ever built

Orbits the Earth to be above the atmosphere (which absorbs X-rays, *luckily!*)

Goes 1/3 of the way to the Moon every 64 hours (2 ½ days)

Chandra takes superbly sharp images: with good spectral resolution too!
X-ray Telescopes are different

Chandra’s mirrors are almost cylinders

X-rays don’t reflect off a normal mirror – they get absorbed. Only by striking a mirror at a glancing angle, about 1°, do X-rays reflect. Then they act like visible light and can be focused.

This makes for loooooooooong telescopes.
The Chandra spacecraft

10 meters (32 ½ ft) from mirror to detector, 1.2 meters (4ft) across mirror

...but focuses X-rays onto a spot only 25 microns across
Chandra detects individual photons

10 meters (32 ½ ft) from mirror to detector, 1.2 meters (4 ft) across mirror

CCD detectors count each X-ray individually

each X-ray knocks free enough electrons to detect as a pulse of electricity

…but can disperse the incoming X-ray light: **Light as Waves**

Delicate gold gratings diffract the light

Chandra provides a great example of how *Quantum wave/particle duality* works in a real machine
Chandra’s sharp focus revolutionizes our understanding

Earth observing satellite equivalents of …

Best X-ray image of whole sky (ROSAT)
Best X-ray images before Chandra (ROSAT)
Chandra images

Any sign of life?
What’s this odd thing?
I get it!
Like looking up the answers at the back of the book

Chandra has solved 20 year old mysteries:

Yes – the background X-ray light is made up of contributions from millions of quasars

No – gas is not pouring down onto the galaxy at the center of a cluster of galaxies. Something stops it, but what?

Yes – Our Milky Way sits in a halo of hot gas stretching to the Andromeda galaxy and beyond

Yes – quasars have hot winds blowing from their cores at 1000 km/s
and many new questions!

2 examples: What *are* we looking at?

- **Antennae – colliding galaxies**
  - Nest of super-bright black holes in binaries – massive progenitors?

- **Centaurus A – nearest quasar**
  - X-ray ‘smoke ring’ from explosion in core?
Chandra’s Revolution through Resolution continues…

Chandra set to run for 5 more years
& may last much longer

All X-ray images are underexposed (limited by photon noise, not systematics)

Deeper looks show

• more morphological detail in diffuse emission,
• more spectral signal-to-noise in each pixel
Merging galaxy Arp 220

• z=0.018 (250 million light years)
• Bolometric luminosity of 10**12 Lsun
• Most energy output in the infrared
• 20-year controversy: star formation or quasar?
• Answer: both, but mostly star formation
Arp 220 nucleus

• Soft extended region of star formation
• Obscured hard point source coincides with radio nuclei
• Off-nuclear point source (ULX binary) $1 \times 10^{39}$ erg/s
Spectra in different regions

- Nuclear spectrum extends to > 7 keV

- Lobe spectrum much softer, < 1 keV
Arp 220: summary

- The AGN in Arp 220 is there!
- but it's only 4E40 erg/s...
- Even correcting for absorption, it can't be the source of most of Arp 220's luminosity
- ULX of 6E39 erg/s at 2.5 kpc from nucleus
- X-ray lobes to 15 kpc each side, correlates with H-alpha, 1E41 erg/s total
- Superwind plus merger remnants?
Cas A (Una Hwang)

- 1 megasecond (11 days)
- Blue: Fe
- Red: Si
- Green: outer shock wave
Cas A

- Si image shows "jets" from asymmetric explosion
C153 (Dan Wang)

• Chandra data shows hot gas stripped from galaxy in cluster
Perseus (Andy Fabian)

- Radio jets make cavities in X-ray cluster gas
- Shocks through cluster due to AGN?
Sgr A*  (Fred Baganoff)

• Flaring from central source: smoking gun for the Galactic supermassive black hole?
Chandra data

- Open competition for Chandra time
- SAO+MIT Chandra X-ray Center (CXC) in Cambridge, MA operates the satellite and supports observers
- CIAO data analysis system (Sun/Linux/OS10)
- chandra.harvard.edu (pretty pictures)
- cxc.harvard.edu (science, calibration, software, proposal submission)
- Our mission: make it possible for non-X-ray specialists to do X-ray astronomy
- All X-ray missions use (pretty much) common data format, good start for Virtual Observatory era
Chandra analysis issues

- Poisson noise - always too few photons
- Position: Astrometry is pretty reliable (1”); PSF degrades when far off-axis
- Time: ACIS has 3s exposure time, HRC can do 16 microsec. Calibration is good.
- Energy: Complicated energy response, varies with detector position and time, and has 'sidelobes'.
- Instrument problems: “CTI” energy resolution variation, “contamination” causing secular sensitivity loss at low energies, bad pixels and columns, background flares
- Universe problems: Cosmic X-ray background, solar X-rays, extinction by interstellar medium, etc.
ACIS analysis

- Multicolor imaging helps in source identification
- Then define source regions and extract instrumental spectra
- Still can't deconvolve instrumental spectral response - forward model fitting required
ACIS analysis

- In detector plane, sources are dithered
- Bad columns can affect effective exposure and introduce spurious time variability
ACIS analysis

- Sensitivity varies by factor 2 across image
- Fairly insensitive to energy
ACIS analysis

- PSF is awesomely good on axis, but (relatively) lousy at the edge of the field (cf 1' vs 16' offaxis)
- PSF is energy dependent
ACIS analysis

- Background flares
- CTI
- Contamination
Chandra

• 5 years of great science: spacecraft operating well
• Combining X-ray and optical (IR, radio) needed to untangle the physics
• High resolution imaging with spectra crucial to separate physical components of sources
• Next call for proposals early 2005